

Newsletter Cosmic Rays / October 2017

October 2017 has been a less active month in the sense of geomagnetic activity. A number of 50 CMEs has been spotted, with 2 CMEs with angular width $90^{\circ} < w < 180^{\circ}$, resulting into distinct modulation of the galactic cosmic rays (source http://sidc.oma.be/cactus/catalog.php). October was one of the quietest months up to now in the sense of proton flux levels of solar flares (SFs). A number of only 2 SFs were spotted, the most energetic one being an M1.1 one on 20/10/2017 at 23:10 UT (start time) from AR2685 with coordinates S12E88 (Fig. 1).

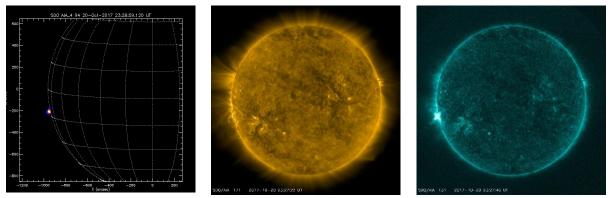
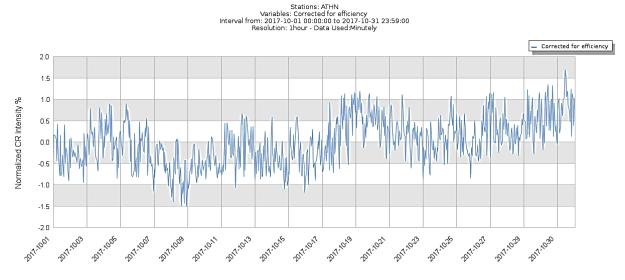


Figure 1: The M1.1 SF of 20/10/2017 at 23:28 UT peak time (from http://www.lmsal.com/solarsoft and http://sdo.gsfc.nasa.gov/data/aiahmi/)

The interaction of high-speed streams of solar wind from coronal holes on October 11-15 and 24-26 as well as disturbed solar wind on October 19, was triggered geomagnetic storms of G1-G2 and G1 levels, respectively. The results of these events were spotted on the cosmic ray intensity as Forbush decreases during this month, recorded at Athens Neutron Monitor Station (cut-off rigidity 8.53 GV) with amplitudes varied from 1% up to almost 2% (Fig. 2).



Data retrieved via NMDB are the property of the individual data providers. These data are free for non commercial use to within the restrictions imposed by the providers. If you use such data for your research or applications, please acknowledge the origin by a sentence like: 'We acknowledge the NMDB database (www.mmdb.eu), founded under the European Union's FP7 programme (contract no. 213007), and the PIs of individual neutron monitors for providing data.'

Figure 2: Hourly corrected for pressure and efficiency values of the cosmic ray intensity recorded by Athens Neutron Monitor Station from 01-31/10/2017 (From the multi station data service of Athens NM Station).

Contact: Prof. H. Mavromichalaki Email: <u>emavromi@phys.uoa.gr</u> http://cosray.phys.uoa.gr